

EE 403/503
Introduction to Plasma Processing

The kinetic Theory of Gases

&

Plasma Properties

Dr. Kasra Etemadi
October 5, 2011

Outline

1- Summary of Kinetic Theory of Gases

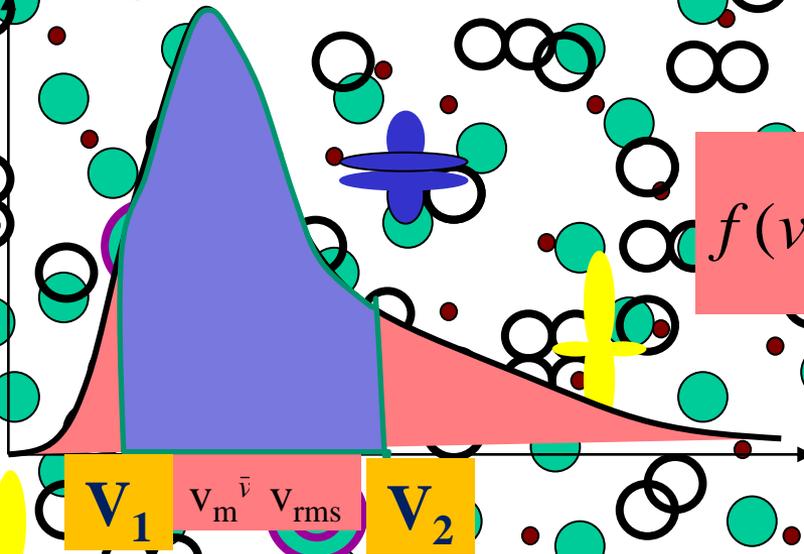
2- Radiation

3- E4

4- Plasma Properties

Plasma

$f(v)$



$$f(v) = 4\pi \left(\frac{m}{2\pi kT} \right)^{3/2} v^2 \exp\left(-\frac{mv^2}{2kT}\right)$$

$$v_m = \left(\frac{2kT}{m} \right)^{1/2}$$

$$\bar{v} = \sqrt{\frac{8kT}{\pi m}}$$

$$v_{rms} = \left(\frac{3kT}{m} \right)^{1/2}$$

E3:

Percentage of particles having speed between V_1 and V_2

$$\frac{N_{V_1 \rightarrow V_2}}{N} = 1 + \frac{2}{\sqrt{\pi}} U e^{-U^2} - \text{erf}(U)$$

The Saha Equation



$$\frac{n_e n_{ion}}{n_a} = \frac{(2\pi m_e kT)^{3/2}}{h^3} \frac{2Z_{ion}}{Z_a} \exp\left(\frac{-e\chi_a}{kT}\right)$$

n_e	number density of electrons	n_{ion}	number density of ions
n_a	number density of atoms	m_e	mass of electron
k	Boltzmann constant	T	Temperature
h	Planck's constant	e	Electron charge
Z_{ion}	Partition function of ion		
Z_a	Partition function of atom		
χ_a	$E_i - \Delta E_i$	$\Delta E_{ionization} = 6.96 \cdot 10^{-13} n_e^{1/3}$ [eV]	
$E_{ionization}$	Ionization potential of neutral atoms		
$\Delta E_{ionization}$	Lowering of ionization potential		

Pressure

$$P_{\text{total}} = p_{\text{molecules}} + p_{\text{atoms}} + p_{\text{ions}} + p_{\text{electrons}}$$

$$p = n k T$$

For multi-temperature Plasmas:

$$p = n_e k T_e + n_{\text{ion}} k T_i + n_a k T_a$$

For equilibrium plasmas:

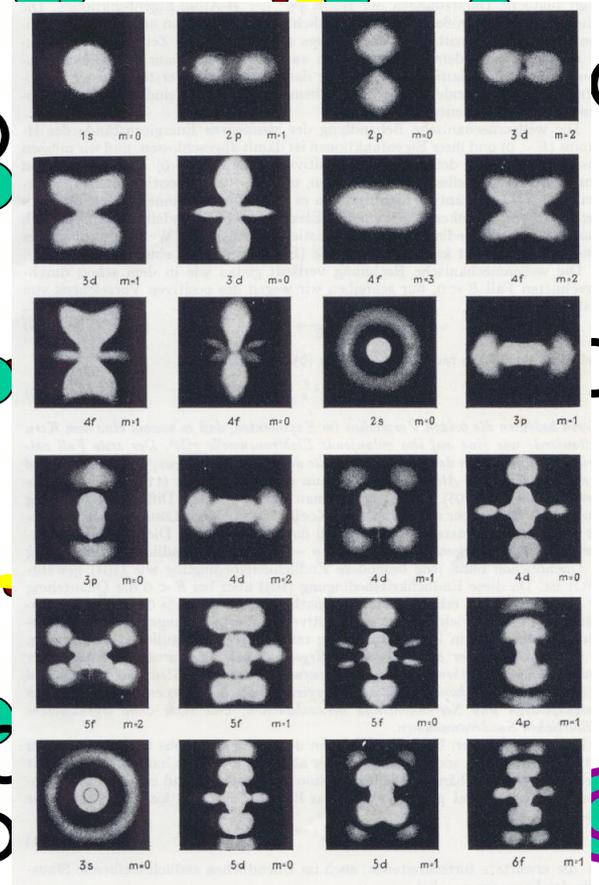
$$p = (n_a + n_{\text{ion}} + n_e) k T$$



Hydrogen

$$n_i = \frac{g_i}{Z_a} n_a \exp\left(\frac{-E_i}{kT}\right)$$

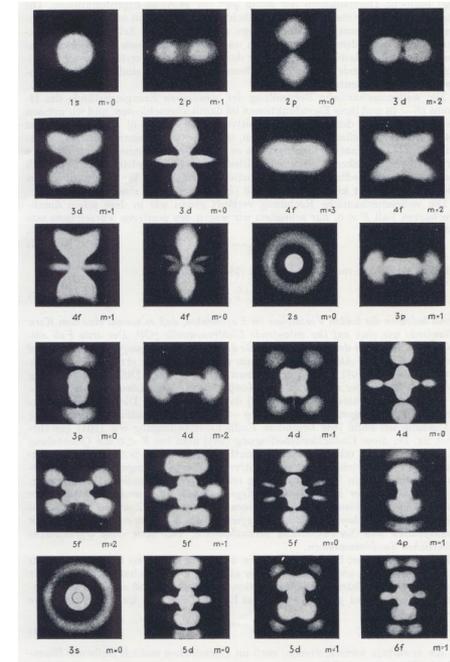
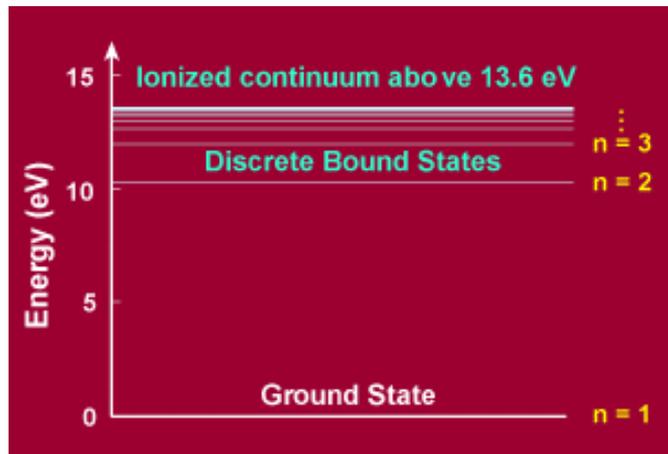
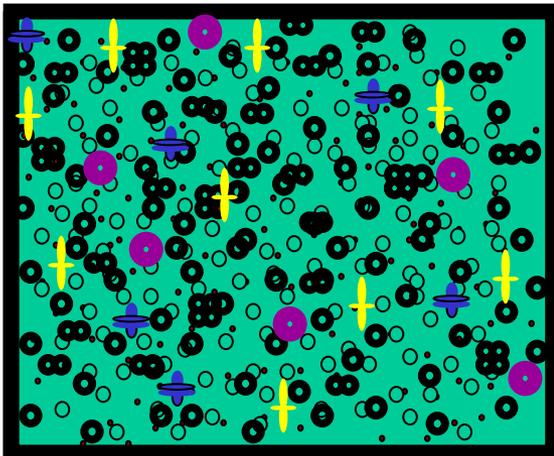
- n_i Number density of atom in i^{th} energy state
- E_i Energy state of the i^{th} energy level
- g_i Degeneracy
- Z_a Partition function
- n_a Total number of atoms
- T Temperature
- k Boltzmann Constant



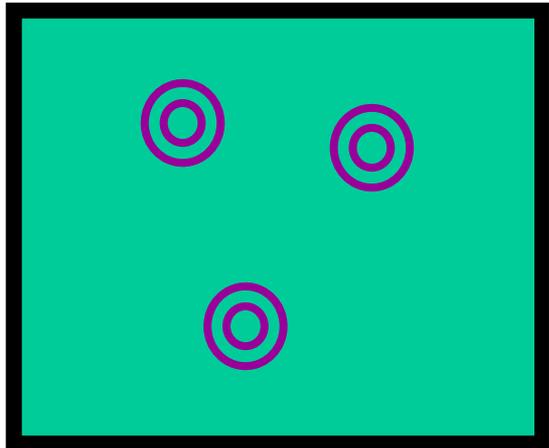
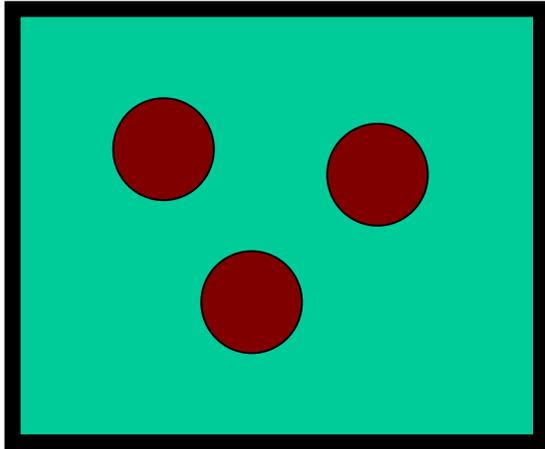
$$n_i = \frac{g_i}{Z_a} n_a \exp\left(\frac{-E_i}{kT}\right)$$

n_i Number density of atom in i^{th} energy state
 E_i Energy state of the i^{th} energy level
 g_i Degeneracy
 Z_a Partition function $Z_a = \sum_{i=1}^{n_{\text{ionization}}} g_i \exp\left(\frac{-E_i}{kT}\right)$
 n_a Total number of atoms
 T Temperature
 k Boltzmann Constant

Hydrogen

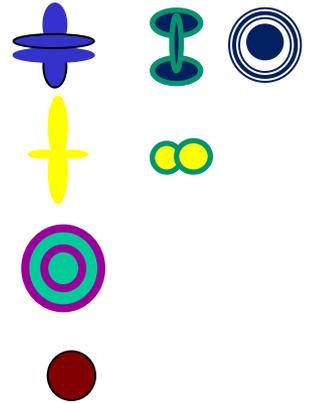


Energy=3E



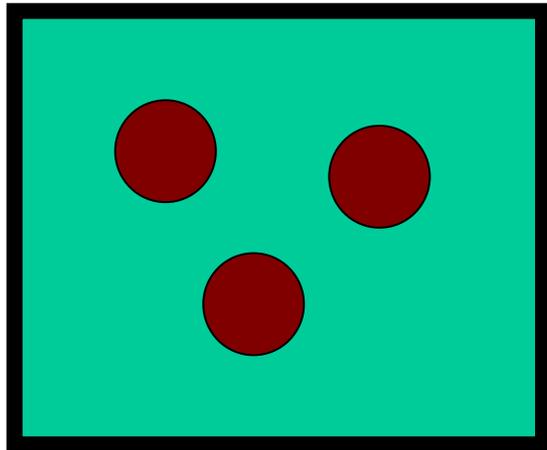
X atom

$E_4=3E$	$n=4$	degeneracy $g_4=3$
$E_3=2E$	$n=3$	$g_3=2$
$E_2=E$	$n=2$	$g_2=1$
$E_1=0$ Ground State	$n=1$	$g_1=1$



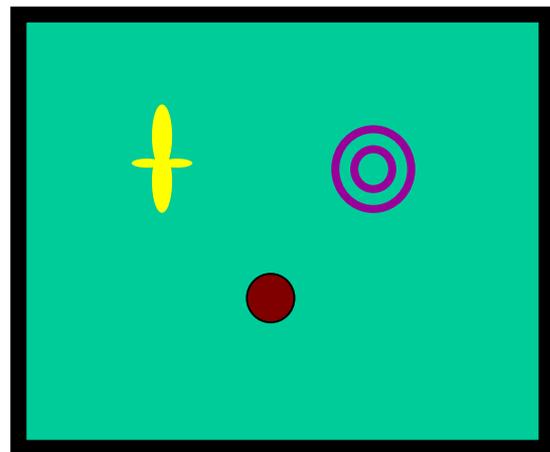
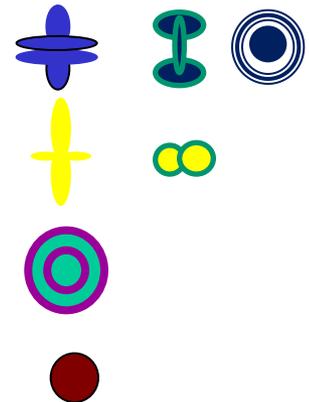
3E1	≡	≡	≡
2E1	==	==	==
E1	●	●	●
0	—	—	—

Energy=3E

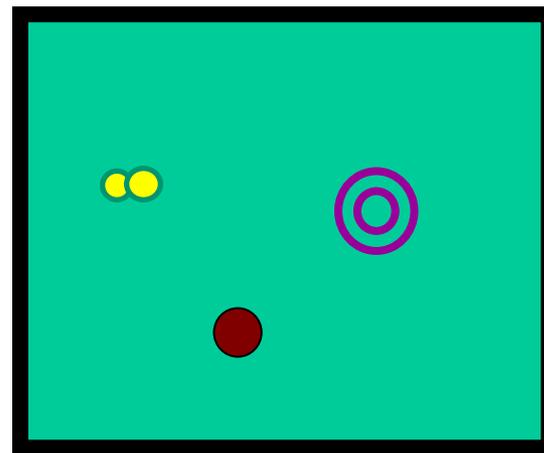


X atom

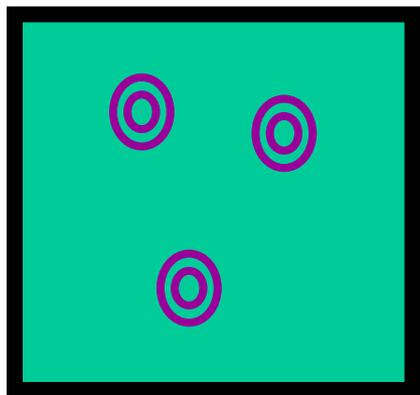
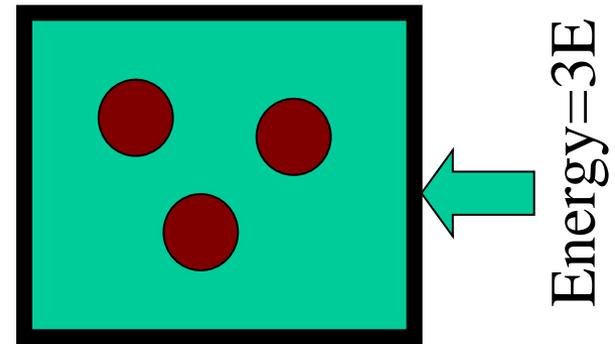
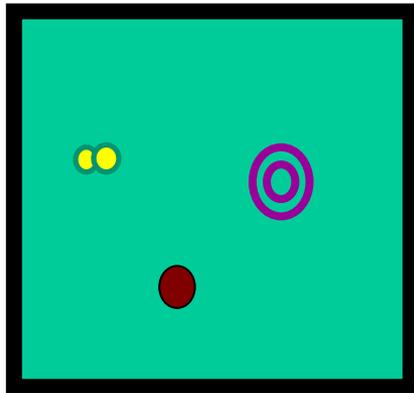
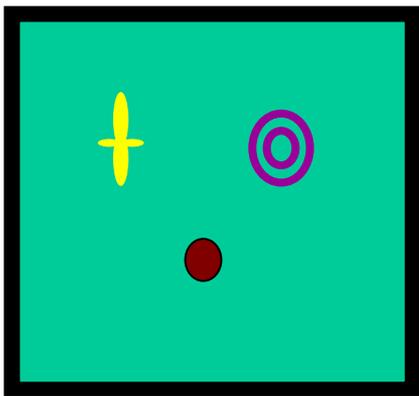
Energy	n	degeneracy
$E_4=3E$	$n=4$	$g_4=3$
$E_3=2E$	$n=3$	$g_3=2$
$E_2=E$	$n=2$	$g_2=1$
$E_1=0$ Ground State	$n=1$	$g_1=1$



3E1	≡	≡	≡
2E1	≡	≡	≡
E1	≡	≡	≡
0	≡	≡	≡



3E1	≡	≡	≡
2E1	≡	≡	≡
E1	≡	≡	≡
0	≡	≡	≡

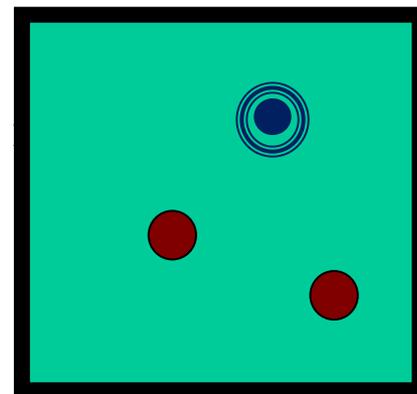
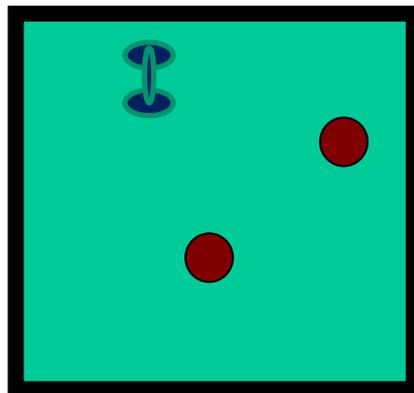
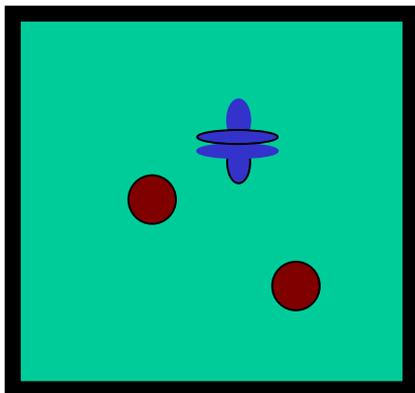


$$n_i = \frac{g_i}{Z_a} n_a \exp\left(\frac{-E_i}{kT}\right)$$

$$Z_a = \sum_{i=1}^{n_{\text{ionization}}} g_i \exp\left(\frac{-E_i}{kT}\right)$$



$E_4=3E$	$n=4$	degeneracy $g_4=3$
$E_3=2E$	$n=3$	$g_3=2$
$E_2=E$	$n=2$	$g_2=1$
$E_1=0$ Ground State	$n=1$	$g_1=1$



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Radiation

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Blackbody Radiation

$$E(\lambda, T) = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$$

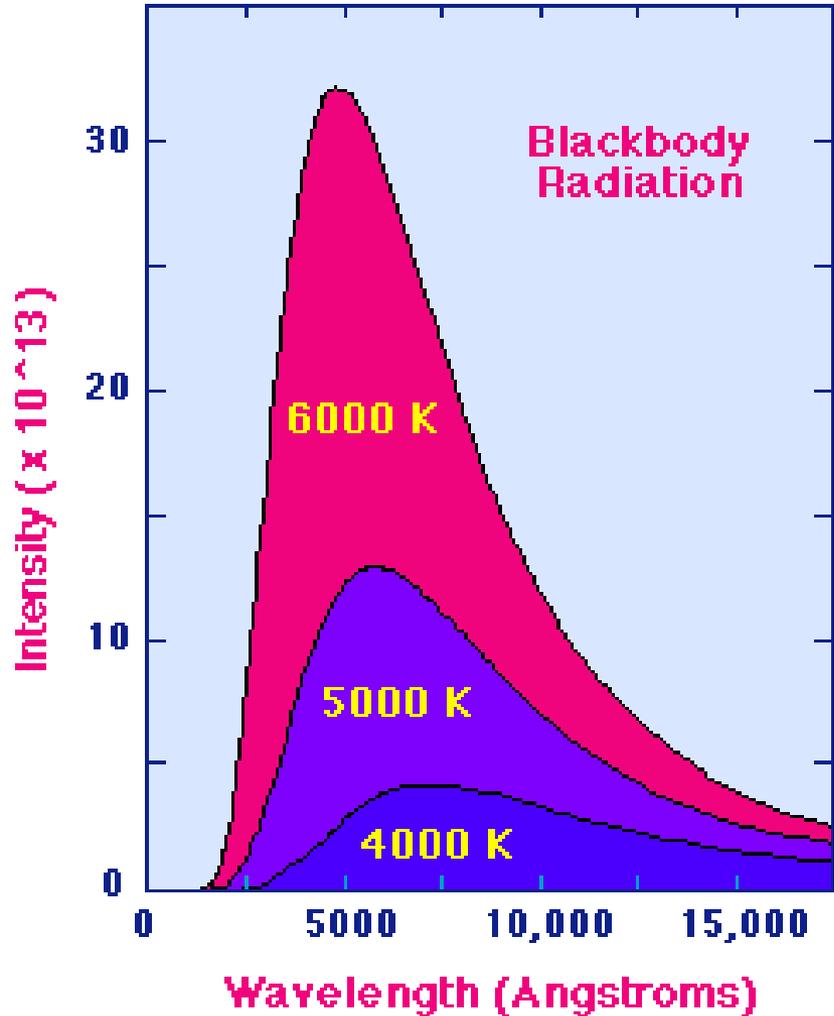
Stefan-Boltzmann Law:

$$E = \sigma T^4$$

$\sigma = 5.67 \times 10^{-8} \quad [Jm^{-2}s^{-1}K^{-4}]$

Wien Displacement Law:

$$\lambda_{Max} = \frac{3 \times 10^7}{T} \quad (\lambda \text{ is in } \text{\AA})$$



Link1

Continuous \longleftrightarrow **Line**

4000 Å

7500 Å



Continuous Spectrum



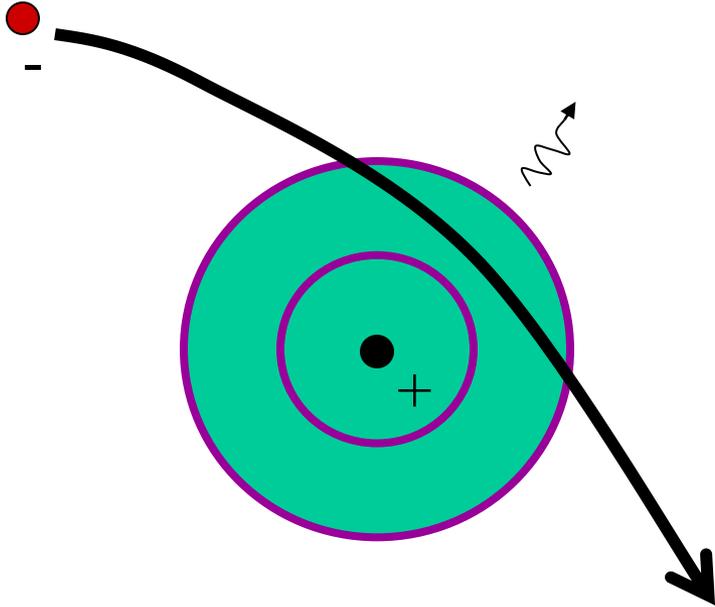
Absorption Spectrum of Hydrogen



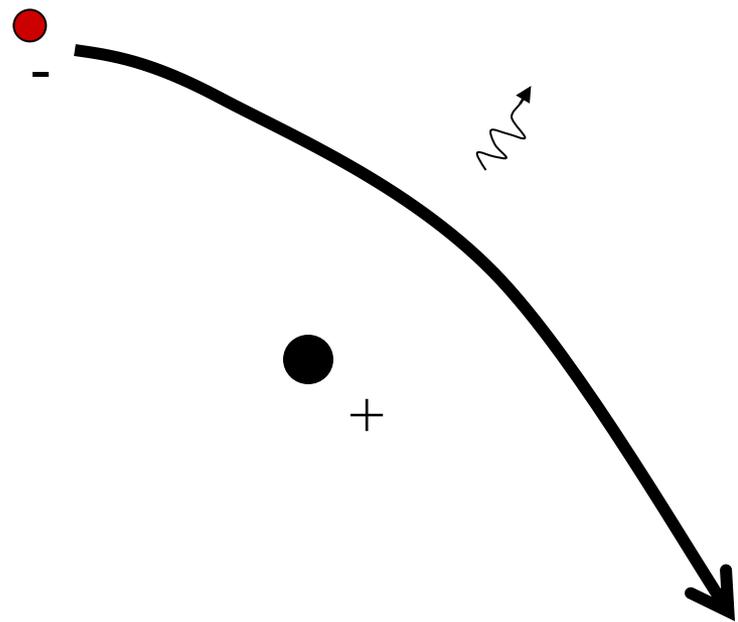
Emission Spectrum of Hydrogen

Continuous Spectrum

electron



electron



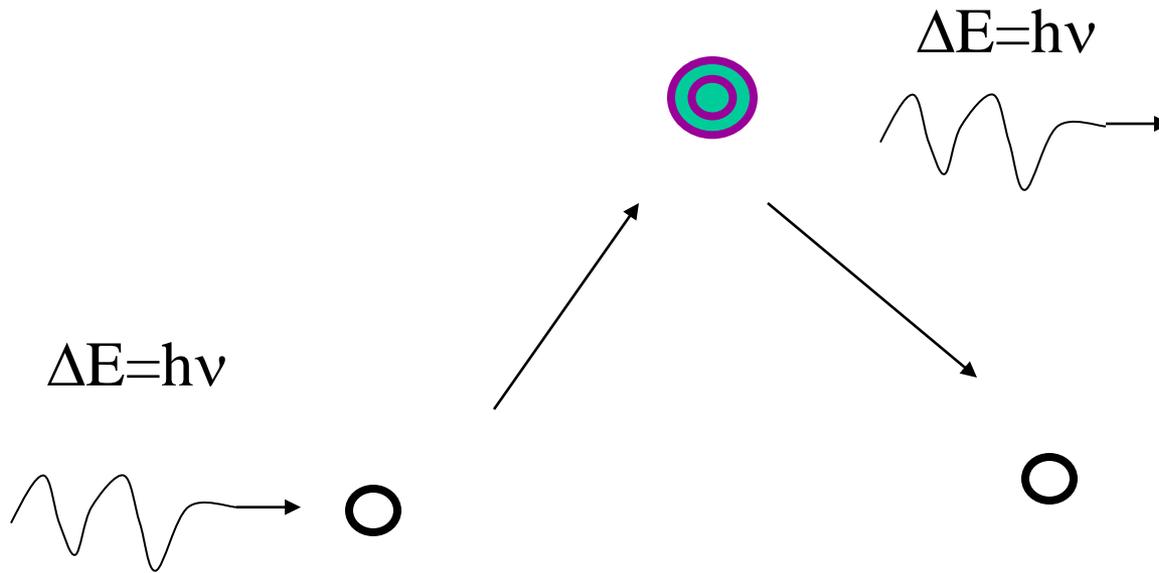
$$\varepsilon_{\lambda}(T) = 1.63075 \times 10^{-28} \frac{N_e^2}{\lambda^2 T^{1/2}} \xi(\lambda, T)$$

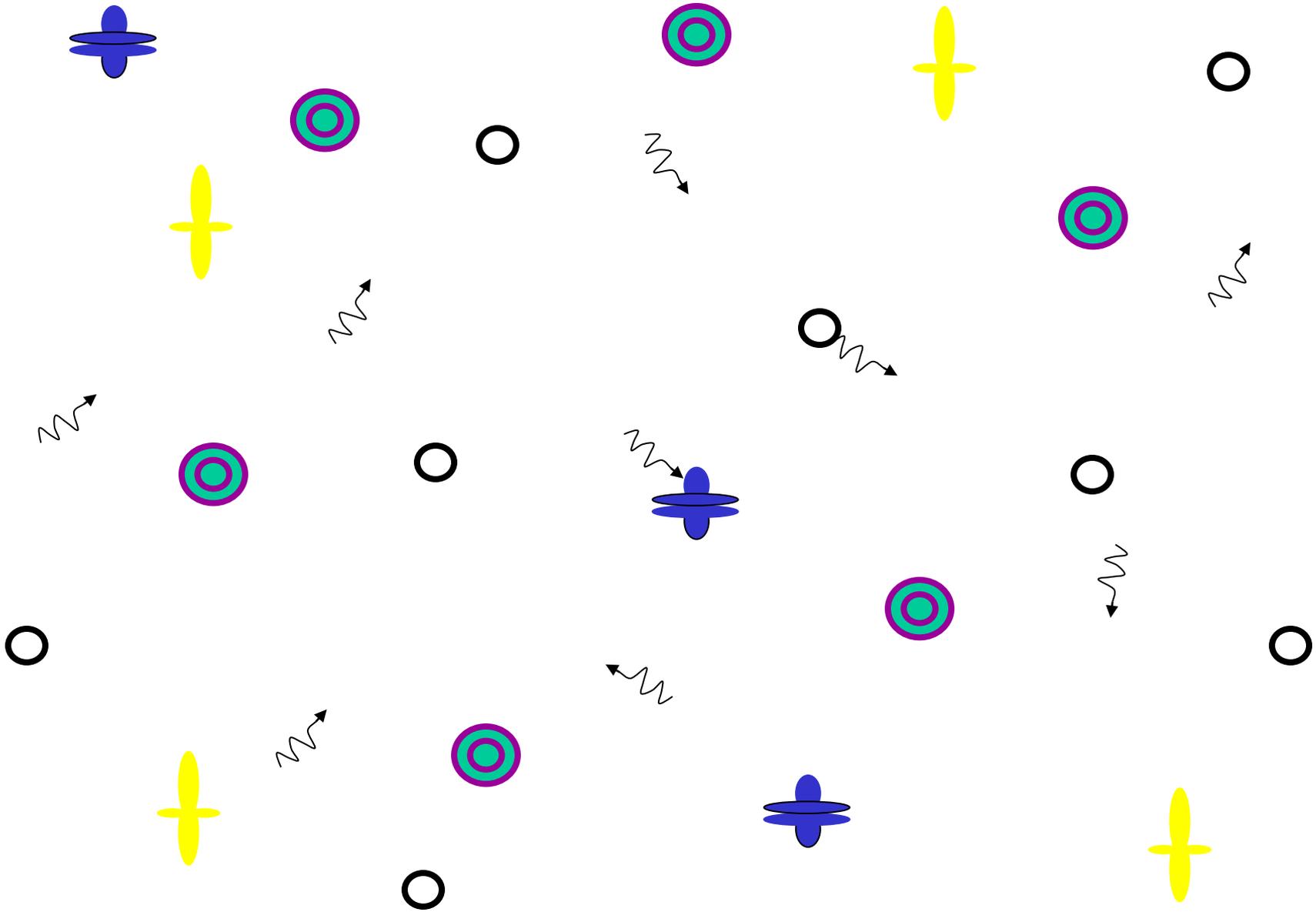
$$\xi(\lambda, T) = \frac{g_i}{Z_i} \xi^{fb}(\lambda, T) (1 - e^{-\frac{c}{\lambda T}}) + \xi^{ff}(\lambda, T) e^{-\frac{c}{\lambda T}}$$

Spectral Lines

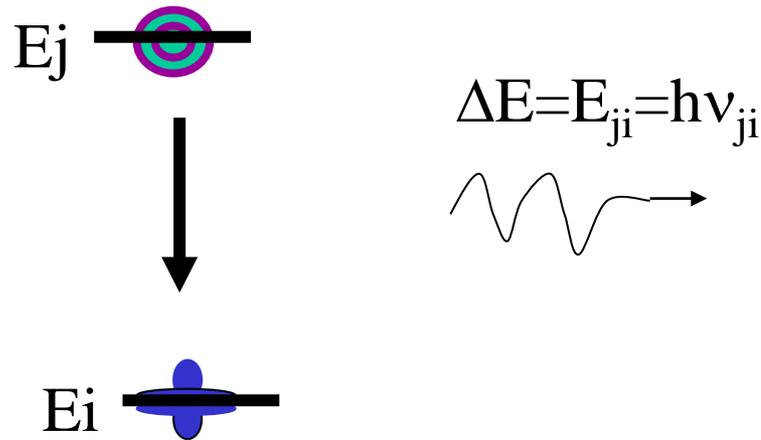
Absorption

Emission





Radiation Energy



$$E_{\lambda} = \frac{1}{4\pi} n_j A_{ji} E_{ji} = \frac{1}{4\pi} n_j A_{ji} h \nu_{ji}$$

$$E_{\lambda} = \frac{1}{4\pi} n_a \frac{g_j}{Z_a} \exp\left(\frac{-E_j}{kT}\right) A_{ij} h \nu_{ji}$$

$$E_{total} = \sum_{\lambda} E_{\lambda} + E_{continuum}$$

Plasmas

Optically Thin



Optically Thick

Maxwellian Distribution

$$f(v) = 4\pi \left(\frac{m}{2\pi kT} \right)^{3/2} v^2 \exp\left(-\frac{mv^2}{2kT}\right)$$

T Maxwellian

Boltzmann Distribution

$$n_i = \frac{g_i}{Z_a} n_a \exp\left(\frac{-E_i}{kT}\right)$$

T Boltzmann

Saha Equation

$$\frac{n_e n_i}{n_a} = \frac{(2\pi m_e kT)^{3/2}}{h^3} \frac{2Z_i}{Z_a} \exp\left(\frac{-e\chi_a}{kT}\right)$$

T Saha

Planck's Function

$$E(\lambda, T) = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$$

T Planck

Complete Thermodynamic Equilibrium

$$T_{\text{Maxwellian}} = T_{\text{Boltzmann}} = T_{\text{Saha}} = T_{\text{Planck}}$$

Local Thermodynamic Equilibrium

$$T_{\text{Maxwellian}} = T_{\text{Boltzmann}} = T_{\text{Saha}}$$

Partial Local Thermodynamic Equilibrium

$$T_{\text{Maxwellian}} = T_{\text{Saha}}$$

Example

For an atmospheric pressure hydrogen plasmas having a temperature of 10,000 K determine the following:

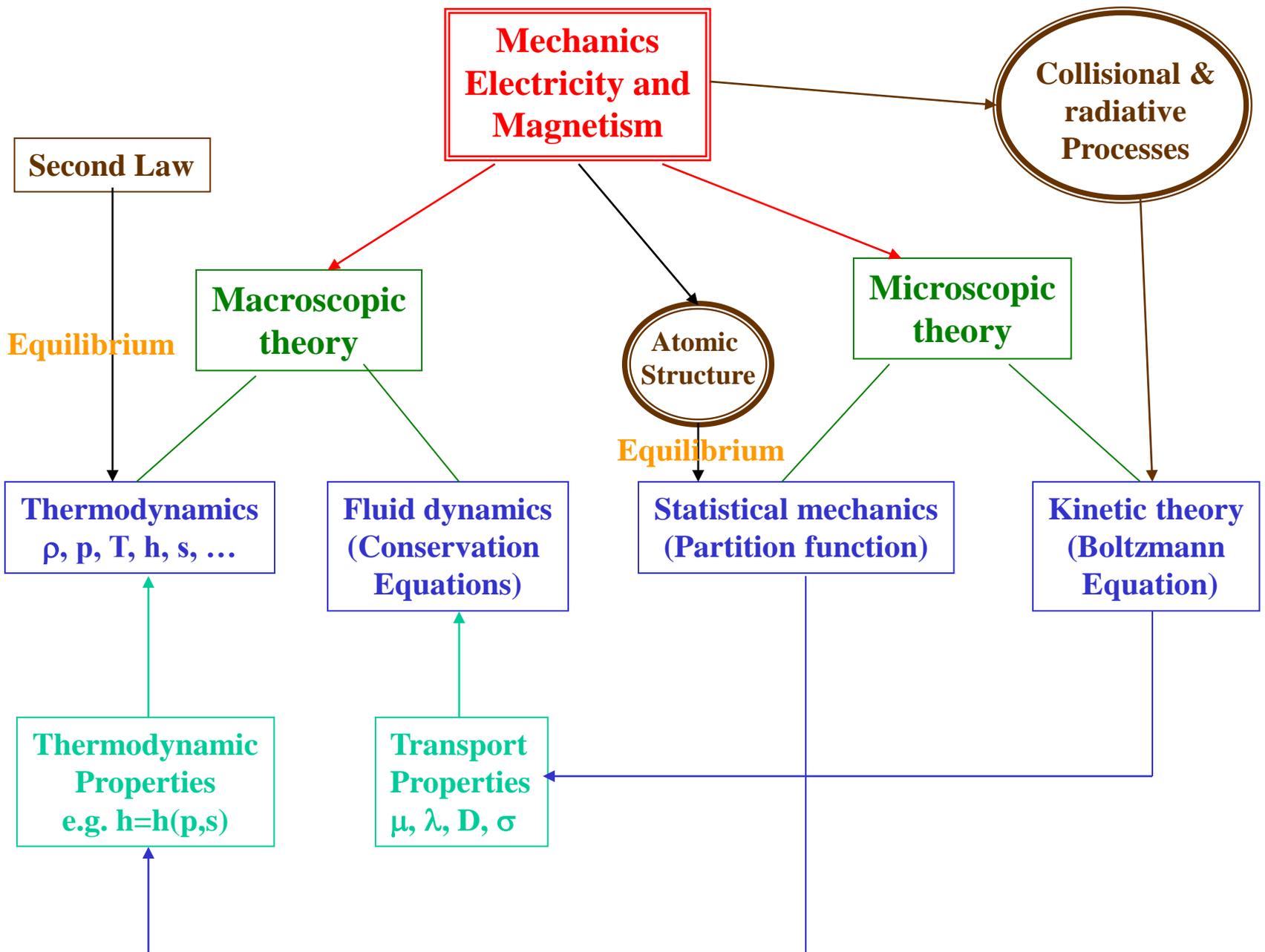
- number density of atoms, ions and electrons
- Maximum number densities of atoms, ions and electrons that can contribute to excitation of hydrogen atoms from ground state.
- Density distribution of energy states of hydrogen atoms
- Intensity [J/cm^3] of hydrogen spectral line at 1215.72 \AA (first excited to ground state, $g_u=8$, $Z_H=2$, $A_{ul}=4.699\text{e}8 \text{ s}^{-1}$)

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Introduction to Plasma Processing

Plasma Properties

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Boltzmann Equation

Net rate of increase of species s , as a result of collisions between particles of species s with those of species r

Species distribution function $f_s(\mathbf{x}, \mathbf{v}, t)$

External forces

$$\frac{\partial}{\partial t} (n_s f_s) + \vec{v} \cdot \nabla n_s f_s + \frac{\mathbf{F}_s}{m_s} \cdot n_s \nabla_v f_s = \sum_r C_{sr}$$

Number of particles n_s

Velocity

Mass of species s

Impact parameter

$$C_{sr} = n_r n_s \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} [f_s(\mathbf{v}'_s) f_r(\mathbf{v}'_r) - f_s(\mathbf{v}_s) f_r(\mathbf{v}_r)] \gamma_{rs} b db dv_r$$

after the collision

Boltzmann Equation

Derivation?

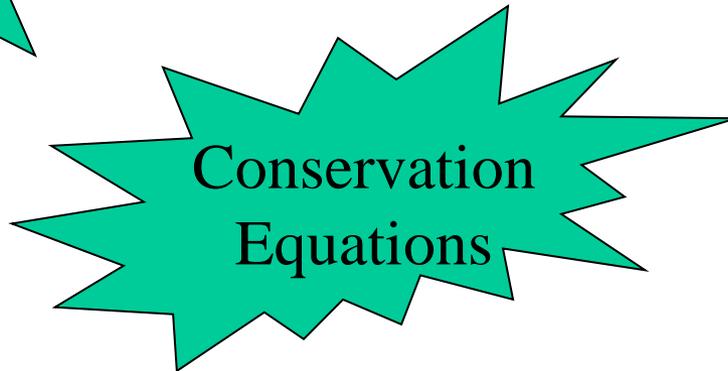
Derivation?

Derivation?

Maxwellian
Distribution

Plasma
properties

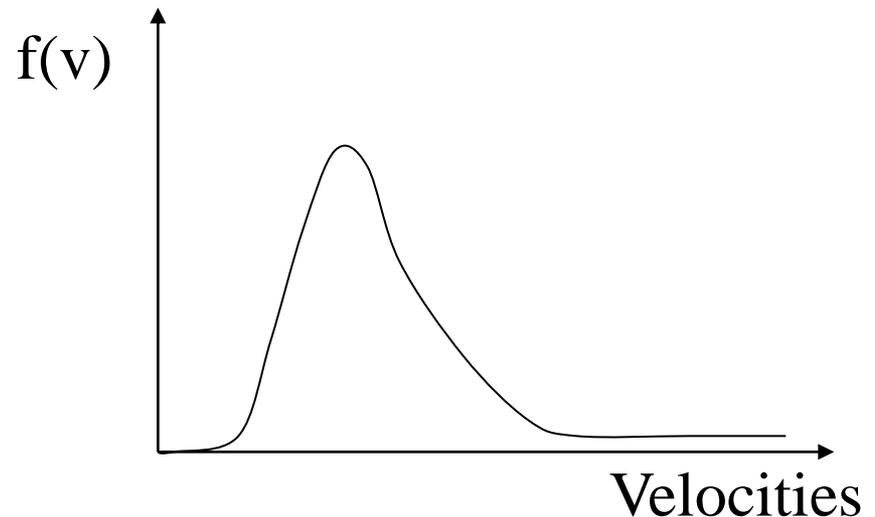
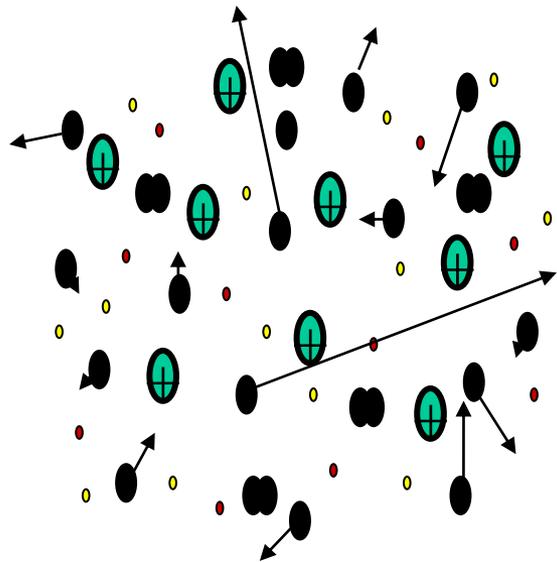
Conservation
Equations



Maxwellian distribution

Neglecting all external forces in Boltzmann equation gives the Maxwell distribution function:

$$f(v) = \frac{dn_v}{dv} = \frac{4n}{\sqrt{\pi}} \left(\frac{m}{2kT}\right)^{3/2} v^2 \exp\left(-\frac{mv^2}{2kT}\right)$$



Conservation Equations

General Form

Dependent
variable

Specific to a particular
Meaning of Φ


$$\frac{\partial}{\partial t} (\rho\Phi) + \nabla \cdot (\rho\vec{u}\Phi) = \nabla \cdot (\Gamma\nabla\Phi) + S$$

Unsteady
term

Convection
term

Diffusion
term

Source
term

Examples:

General:

$$\frac{\partial}{\partial t}(\rho\Phi) + \nabla \cdot (\rho \vec{u}\Phi) = \nabla \cdot (\Gamma \nabla \Phi) + S$$

$$\Phi = 1 \quad \frac{\partial}{\partial t}(\rho) + \nabla \cdot (\rho \vec{u}) = 0$$

$$\Phi = \vec{u} \quad \frac{\partial}{\partial t}(\rho \vec{u}) + \nabla \cdot (\rho \vec{u}\vec{u}) = \nabla \cdot (\Gamma \nabla \vec{u}) + \nabla p + S_m$$

$$\Phi = m_l \quad \frac{\partial}{\partial t}(\rho m_l) + \nabla \cdot (\rho \vec{u} m_l) = \nabla \cdot (\Gamma \nabla m_l) + R_l$$

$$\Phi = h \quad \frac{\partial}{\partial t}(\rho h) + \nabla \cdot (\rho \vec{u} h) = \nabla \cdot \left(\frac{k}{c_p} \nabla \Phi \right) + S_h$$

To utilize the conservation equations in your application, you need:

-a set of conservation equations that are required to describes your applications

-Set up your initial condition and boundary conditions

-Provide the required thermodynamic and transport properties

Solve the conservation equations analytically or numerically

Transport Properties

1- Diffusivity (Page 49)

Movement of gas particles from high density to low density gas by random walk

Particle Flux $\Gamma = n v_d = -D \nabla n$ Diffusion Coefficient

Density n Drift Velocity v_d

Fick's Law

Average Velocity

$$D \approx \frac{1}{3} v_c \lambda^2 = \frac{1}{3} v \lambda = \frac{1}{3} v^2 \tau$$

Collision Frequency v_c Mean Free Path λ

if $\frac{\lambda}{n} \frac{\partial n}{\partial x} = \frac{\lambda}{L} \ll 1$

For a monolithic gas:

$$D = 2.06 \times 10^{-12} \frac{\sqrt{mT}}{\sigma}$$

σ

Cross section for binary hard-sphere collisions

2- Electrical Conductivity

Current Density $\rightarrow J = \sigma E$ [A/m²]

Electrical Conductivity $\rightarrow \sigma = \frac{e^2 n_e}{m_e v_e}$

3- Thermal Conductivity

Heat Flux $\rightarrow q = -k \nabla T$

Thermal Conductivity $\rightarrow \kappa \approx f \bar{v} \lambda$

For a monolithic gas:

$$\kappa = \frac{25}{64} \frac{fk}{\sigma} \left(\frac{\pi k T}{m} \right)^{1/2} = \frac{C_v \bar{v}}{3 N_{Avag} \sigma_{cross}}$$

Detailed description: This block contains the derivation of thermal conductivity for a monolithic gas. It starts with the general formula $\kappa \approx f \bar{v} \lambda$. The term f is labeled 'Degree of Freedom' with a green arrow pointing to the fk term in the subsequent equation. The term \bar{v} is labeled 'Heat Capacity' with a green arrow pointing to the $C_v \bar{v}$ term in the final equation. The final equation shows the derivation: $\kappa = \frac{25}{64} \frac{fk}{\sigma} \left(\frac{\pi k T}{m} \right)^{1/2} = \frac{C_v \bar{v}}{3 N_{Avag} \sigma_{cross}}$. The term N_{Avag} is the Avogadro number, and σ_{cross} is the cross-sectional area.

4- Viscosity

Force resulting from the net transport of momentum from one region to another

$$F = \eta \frac{du}{dz} \quad [\text{N/m}^2]$$

Coefficient of Viscosity

$$\eta = \frac{1}{3} \frac{m \bar{v}}{\sigma_{cross}} \quad [\text{Ns/m}^2]$$

Cross section for the binary hard-sphere collision

5- Mobility

Drift Velocity

$$U_{e_d} = -\overset{\text{Mobility}}{\downarrow} \mu_e \overset{\text{Electric Field}}{\leftarrow} E$$

$$\mu_e = \frac{3}{4} \frac{e}{n \sigma} \left(\frac{2\pi}{m_e kT} \right)^{1/2}$$

Thermodynamic Properties

[Link1](#)

1- Internal Energy

2- Enthalpy

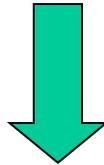
3- Entropy

4- Gibbs Free Energy

Partition Function

$$Z = \sum_s g_s e^{-\frac{E_s}{kT}}$$

$$E_s = E_{\text{translational}} + E_{\text{electronic}} + E_{\text{vibrational}} + E_{\text{rotational}} + E_{\text{ionization}}$$



$$Z = Z_{\text{translational}} \times Z_{\text{electronic}} \times Z_{\text{vibrational}} \times Z_{\text{rotational}} \times Z_{\text{ionization}}$$

$$Z = \sum_s g_s e^{-\frac{E_s}{kT}}$$

$$Z = Z_{\text{translational}} + Z_{\text{electronic}} + Z_{\text{vibrational}} + Z_{\text{rotational}} + Z_{\text{ionization}}$$

$$Z_{\text{translational}} = \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \frac{V}{h^3} e^{-\frac{mv^2}{2kT}} d\vec{v} = \frac{V}{h^3} (2\pi mkT)^{3/2}$$

$$Z_{\text{electronic}} = \sum_i (2j+1) e^{-\frac{E_i}{kT}} = \sum_i 2n^2 e^{-\frac{E_i}{kT}}$$

For Hydrogen

$$Z_{\text{ionization}} = e^{-\frac{\chi_i}{kT}}$$

For molecule made of two identical atoms (rigid rotator):

$$Z_{rotational} = \frac{1}{2} \sum_{j=0}^{\infty} (2j+1) e^{-\left[\frac{h^2}{8\pi^2\Theta} j(j+1) \right] / kT} = \frac{8\pi^2\Theta}{2h^2} kT$$

Approximate Solution

For harmonic oscillation

$$Z_{oscillation} = \sum_{\nu=0}^{\infty} e^{-\frac{h\nu(\nu+1/2)}{kT}} = \frac{1}{2 \sinh(h\nu / 2kT)}$$

$$Z = \sum_s g_s e^{-\frac{E_s}{kT}}$$

Total partition function for a system composed of N identical, Indistinguishable particles:

$$Z_{total} = \frac{Z^N}{N!}$$

$$U = \frac{kT^2}{N} \frac{\partial}{\partial t} (\ln Z_{total})$$

Internal Energy

Sum of energies of individual atoms or molecules

$$S = \frac{k}{N} \ln Z_{total} + \frac{U}{T}$$

Entropy

Definition: $dS = \frac{dQ}{T}$

$$S = k \ln W$$

Probability of system

$$H = U + P \frac{V}{N}$$

Enthalpy

Definition: $H=U+pV$

$$G = H - TS$$

Gibbs Free Energy

Definition: $G=U-TS+pV$

Equilibrium Concept

In a Complete Thermal Equilibrium (CTE) system the following conditions are met.

- 1- Radiation emitted by the system follow Black Body Radiation
- 2- All species have Maxwellian distribution
- 3- Kinetic equilibrium must exist, ie., $T_e = T_h$ (E/P is small, T is high)
- 4- Collision process establish excitation and ionization equilibrium
- 5- Spatial variations are small

Local Thermodynamic Equilibrium (LTE) means that conditions 2-5 are met.

Partial Local Thermodynamic Equilibrium (PLTE) means that one condition is only partially validated (e.g. excitation may not exactly follow the Boltzmann distribution).