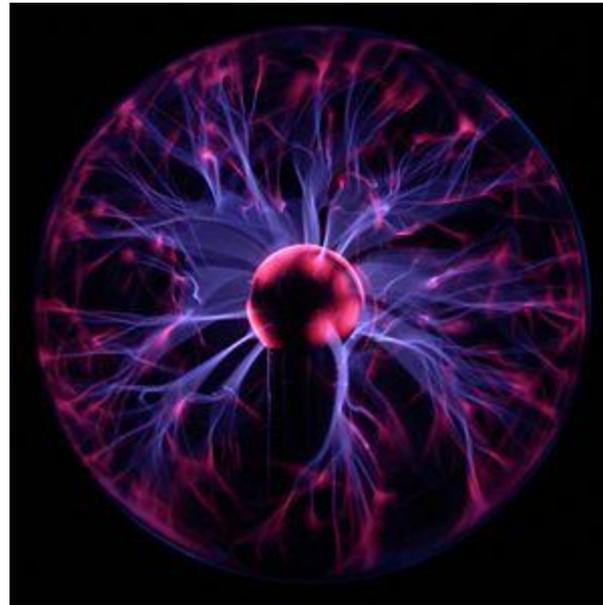


Basic Atomic Theory



EE 403/503

Introduction to Plasma Processing

September 7, 2011

Outline

1- Summary

2- Introduction to Atomic Theory

Bohr's Atomic Model

Quantum Mechanical Model

3- Some Definition

Temperature

Pressure

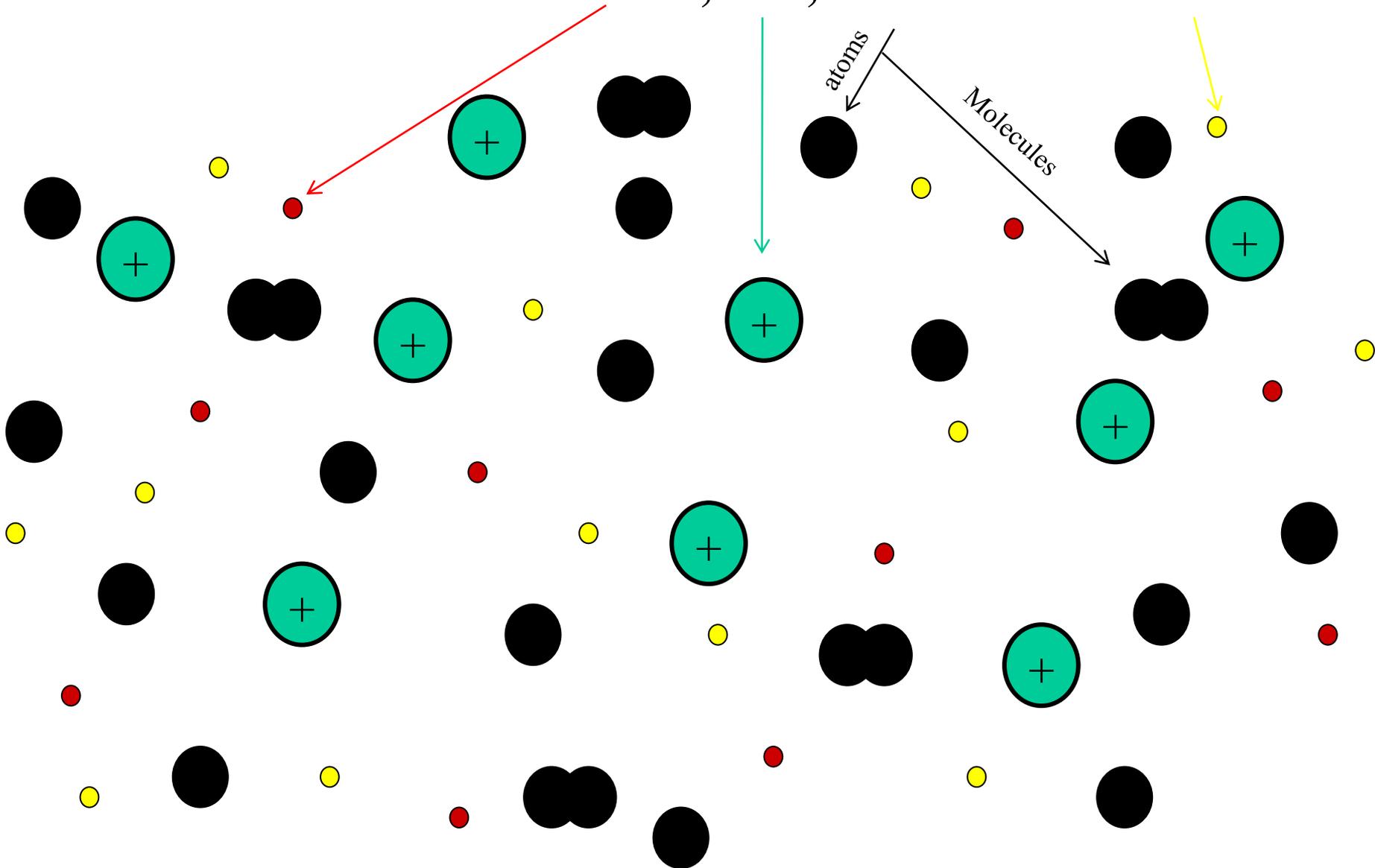
3- Projects

Website

Subject Areas

Summary

Plasma is a Mixture of electrons, ions, neutrals and Photons



Plasma Species

(Structures and Properties)

Electrons:

mass= 9.11×10^{-31} Kg, Charge= 1.602×10^{-19} C and
 $F=e(E+v \times B)$ Lorentz Force

Photons:

$E=h\nu$, $c=\lambda\nu$ and $p=h/\lambda$

$h=6.626 \times 10^{-34}$ Js $c=3 \times 10^8$ m/s

Neutrals:

Atoms & molecules have internal structure

Therefore, they have **kinetic** + **internal** energies

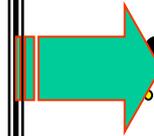
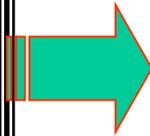
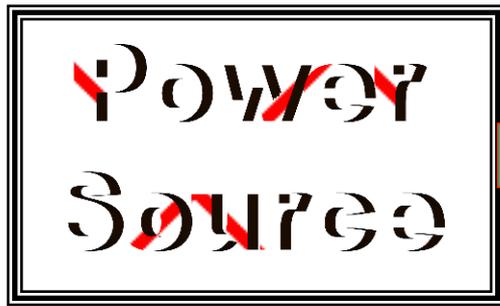
Ions: Almost the same mass as neutrals

Positive (H^+ & Cl_2^+) and negative ions (O_2^- & SF_6^-)



[Omega Centauri](#) from Hubble Telescope

Plasma Generation



Coupling

Breakdown

Maintenance

Power Supply:

Power (currents & Voltages)
Impedance
Frequency

Beams:

Laser
Electrons
ions

Materials:

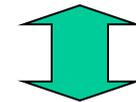
Solid
Liquid
Gas

Plasma Container:

Geometry
Pressure
contaminations

Plasmas:

Concentrations
Velocity distributions
(or temperatures?)



Properties

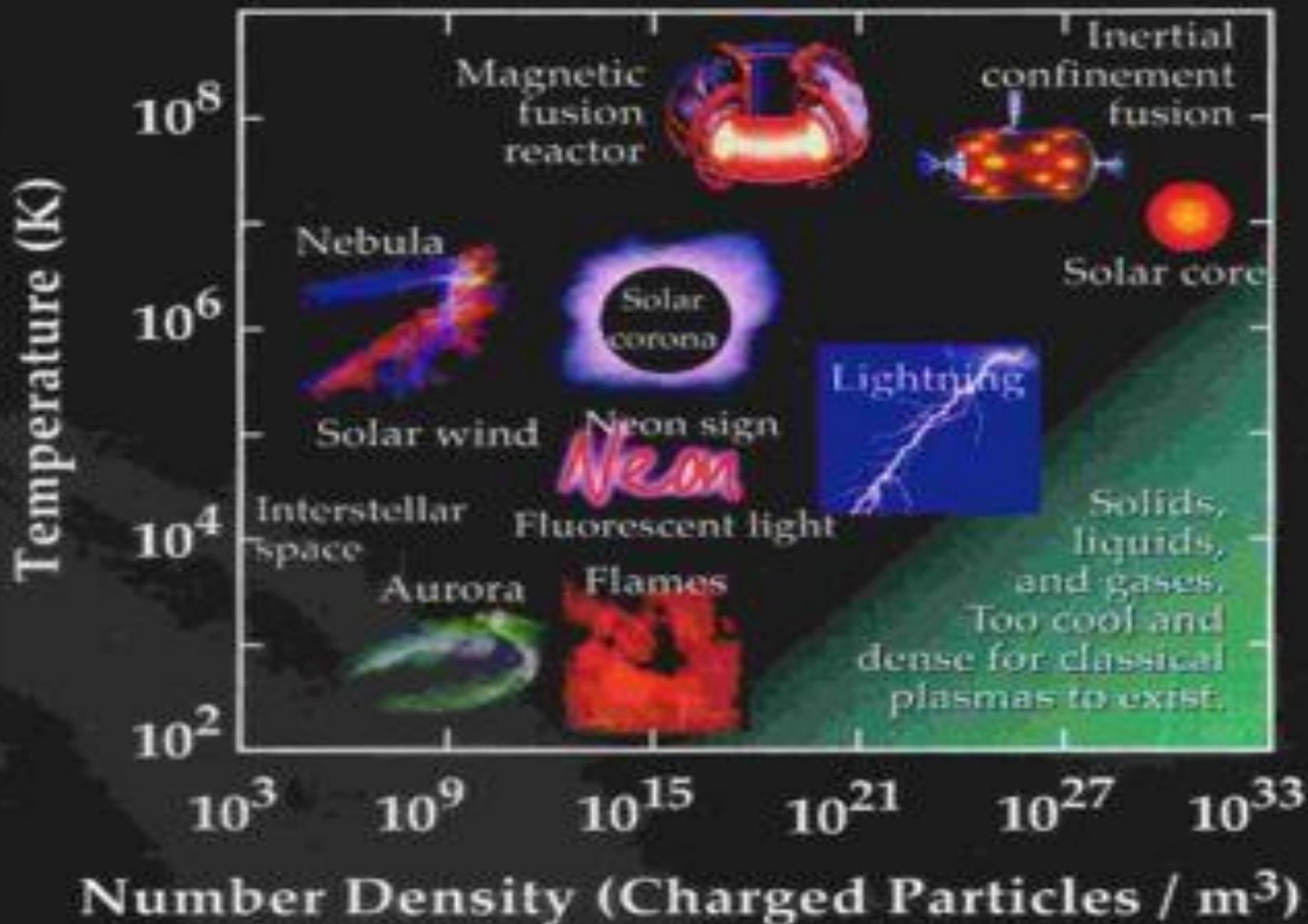
Thermodynamic:

enthalpy, entropy, internal energy, free energy, ...

Transport:

viscosity, electrical and thermal conductivity, mobility, ...

Plasmas - The 4th State of Matter

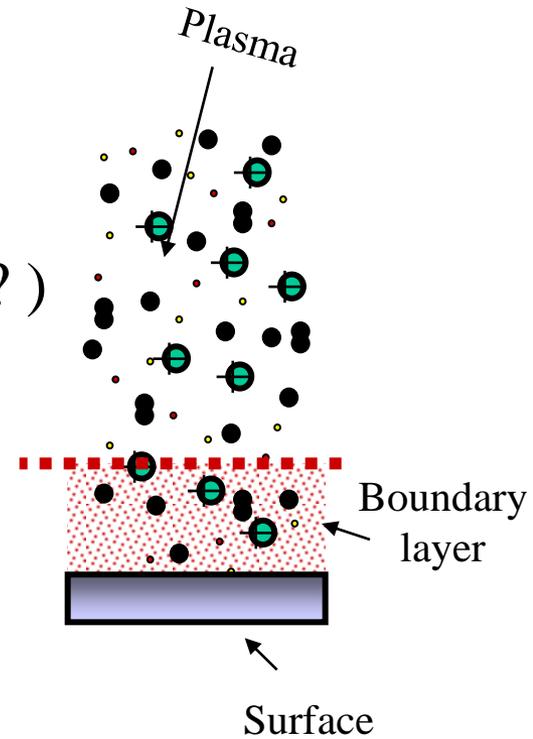


Understanding of plasma and plasma-surface interaction is essential in working with and improving plasma processing applications.

Plasma

- Species concentrations
- Velocity distributions of species (or temperatures?)

Species structure and properties
Interactions among species
Species response to external fields (electric and magnetic)



Plasma-Surface Interactions

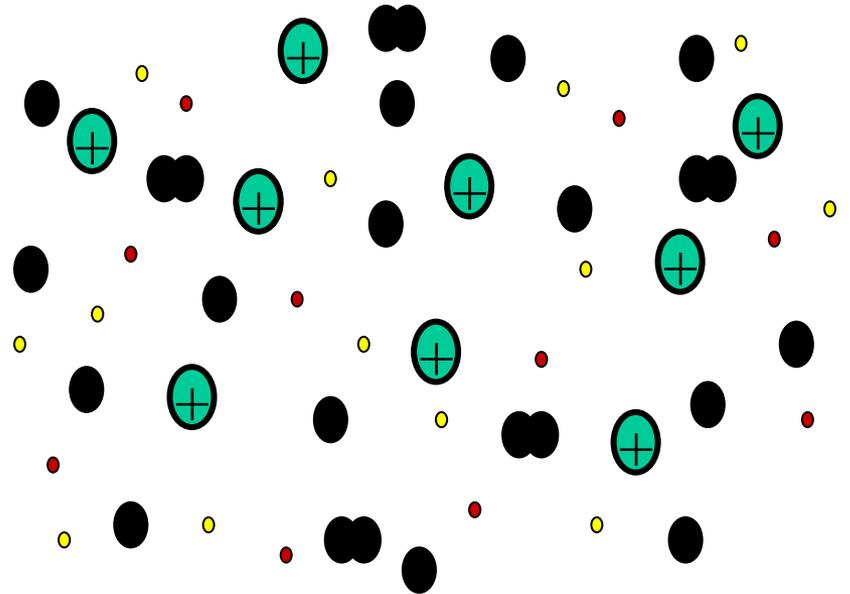
Variations of species concentrations and velocity distributions in the boundary layer between plasma and surfaces

The order of our studies:

- 1- Species structure and properties
- 2- Interactions among species
- 3- Species reactions to external fields (electric and magnetic)
- 4- Species temperatures (or velocity distributions)
- 5- Species concentrations
- 6- Plasma-Surface Interaction
- 7- Plasma Characteristics
- 8- Plasma Discharges
- 9- Plasma Applications

What is Temperature in an ideal gas?

$$E = \frac{3}{2}kT = \frac{1}{2}m\overline{v^2}$$



$$T = f(v)$$

E1-Example 1:

How fast is an Electron and a Hydrogen atom at room temperature?

$$\frac{3}{2}kT = \frac{1}{2}mv^2$$

$$v = \sqrt{\frac{3kT}{m}}$$

$$v = 1.2 \times 10^5 \text{ m/s}$$

Hydrogen:

$$m_{\text{H}} = 1.66 \times 10^{-27} \text{ Kg}$$

$$k = 1.38 \times 10^{-23} \text{ J/K (Boltzmann's Constant)}$$

$$T = 293\text{K}$$

$$v_{rms} = \sqrt{\frac{3kT}{m}}$$

$$v_{rms} = 2703 \text{ m/s}$$

Problem 1



A helium non-thermal plasma used in biomedical applications, the gas temperatures are close to room temperatures. However, the electrons have average energies around 3 eV. Find the average velocity of electrons, He and He⁺.

$$m_{\text{He}} = 6.65 \times 10^{-27} \text{ Kg}; \quad \bar{v} = 0.9213 \times v_{rms}$$

$$E = \frac{3}{2}kT = \frac{1}{2}m\overline{v^2}$$

$$v_{rms,He} = \sqrt{\frac{2E_{He}}{m_{He}}} = \sqrt{\frac{3kT_{He}}{m_{He}}} = \sqrt{\frac{3 \times 1.38 \times 10^{-23} \times 293}{6.65 \times 10^{-27}}}$$

$$v_{rms,He} = 1.35 \times 10^3 \quad m/s \Rightarrow \bar{v}_{He} = 0.9213 \times v_{rms,He} = 1.24 \times 10^3$$

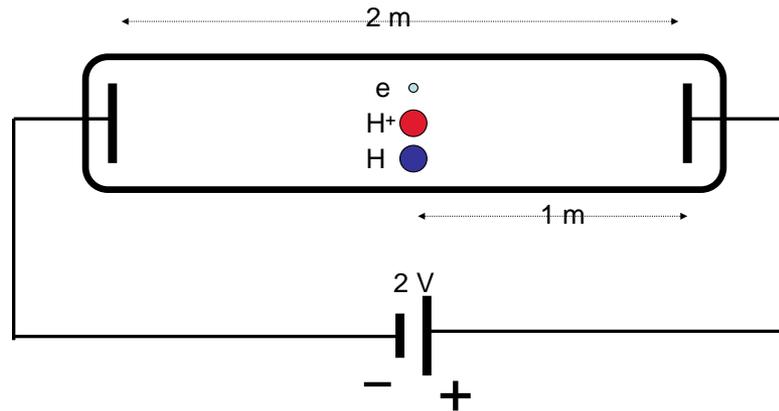
$$v_{rms,He^+} = v_{rms,He} \quad m/s$$

$$v_{rms,e} = \sqrt{\frac{2E_e}{m_e}} = \sqrt{\frac{2 \times 3 \times 1.6021 \times 10^{-19}}{9.1 \times 10^{-31}}} = 1.03 \times 10^6 \quad m/s$$

$$\bar{v}_e = 9.47 \times 10^5 \quad m/s$$

E1: Example 2

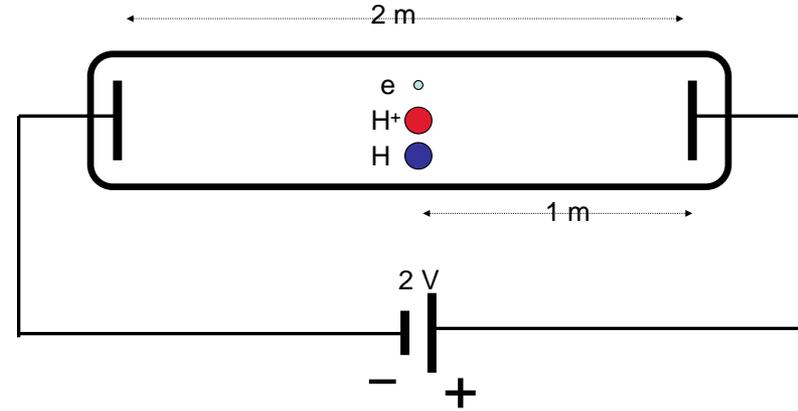
In the vacuum tube shown in the figure below, determine the location, velocity and energy of e, H and H⁺ at the time when the electron arrives at one of the electrodes. All three particles are initially at rest.



$$\begin{aligned}
 & F = ma \\
 & \Downarrow \\
 & a = \frac{F}{m} \qquad F = qE \qquad V = E \times d \Rightarrow E = \frac{V}{d} \\
 & x = \frac{1}{2} at^2 = \frac{1}{2} \left(\frac{F}{m} \right) t^2 = \frac{1}{2} \frac{qE}{m} t^2 = \frac{1}{2} \frac{q \left(\frac{V}{d} \right)}{m} t^2 = \frac{qV}{2md} t^2
 \end{aligned}$$

}
}

Determine the location, velocity and energy of e, H and H⁺ at the time when the electron arrives at one of the electrodes. All three particles are initially at rest.



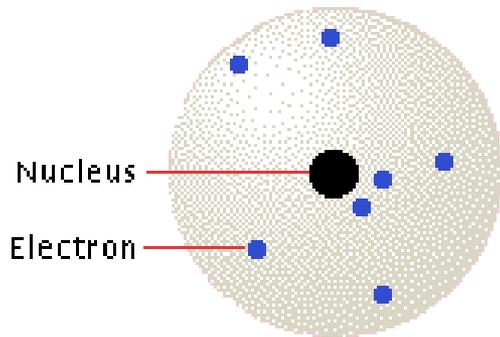
$$x = \frac{1}{2} at^2 = \frac{qV}{2md} t^2$$

$$x_e = \frac{1.6 \times 10^{-19} \times 2}{2 \times 9.11 \times 10^{-31} \times 2} t_e^2 = 1 \quad \Rightarrow \quad t_e = 3.4 \times 10^{-6}$$

$$x_{H^+} = \frac{1.6 \times 10^{-19} \times 2}{2 \times 1.66 \times 10^{-27} \times 2} \times (3.4 \times 10^{-6})^2 = 5.57 \times 10^{-4}$$

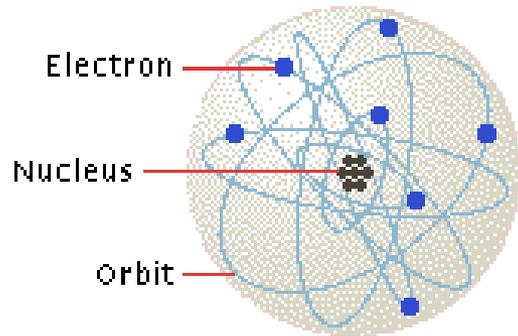
$$x_H = 0$$

Structure of an Atom



Nucleus
Electron

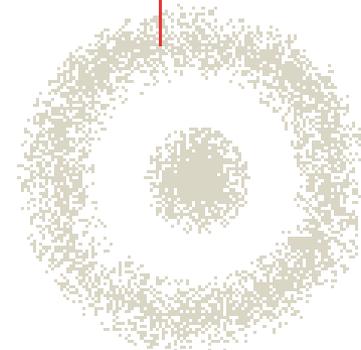
The Rutherford Model
pictured the atom as a miniature solar system with the electrons moving like planets around the nucleus. (1871-1937)



Electron
Nucleus
Orbit

The Bohr Model
(1861-1962) defined the orbits in order to explain the stability of the atom.

The s orbital: Electrons with no angular momentum occupy regions of space like this. Shading shows probability of finding an electron at that distance.



The Schrödinger Model
abandoned the idea of precise orbits, replacing them with a description of the regions of space (called orbitals) where the electrons were most likely to be found. (1887-1961)

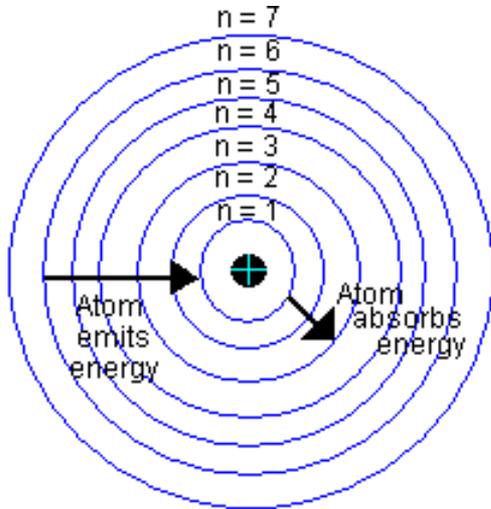
[The Structure of Atoms](#)

(6:11 Min)

[Quantum Mechanics:
The Uncertainty Principle](#)

(5.48 Min)

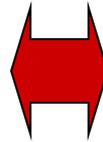
Bohr's Theory



$$\Delta E = R_H \left[\frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$$

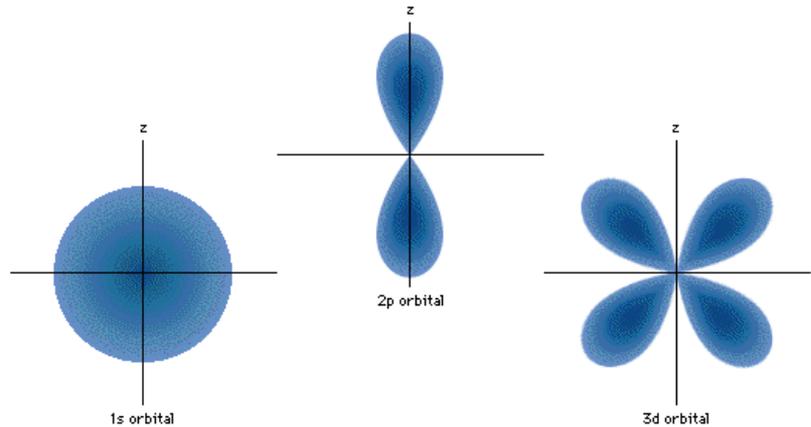
Rydberg constant

$$E_n = 13.5992 \left(1 - \frac{1}{n^2} \right) \quad [eV]$$



Quantum Mechanics

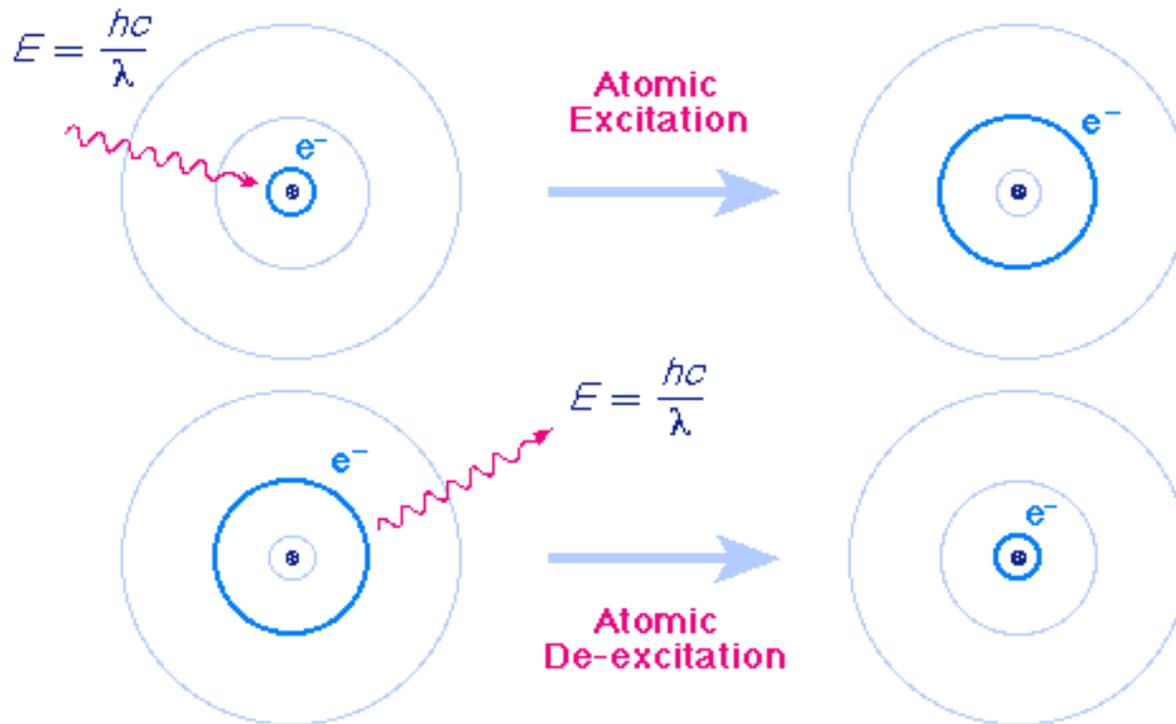
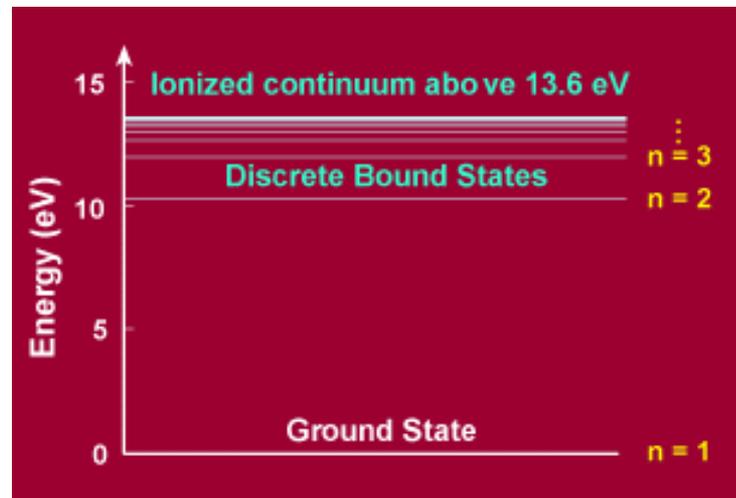
(Schroedinger Equation)



$$i\hbar \frac{\partial \psi(\mathbf{r}, t)}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi(\mathbf{r}, t) + V(\mathbf{r})\psi(\mathbf{r}, t)$$

$$\langle A \rangle = \int \psi^*(\mathbf{r}, t) \hat{A} \psi(\mathbf{r}, t) = \int \psi^*(\mathbf{r}) \hat{A} \psi(\mathbf{r})$$

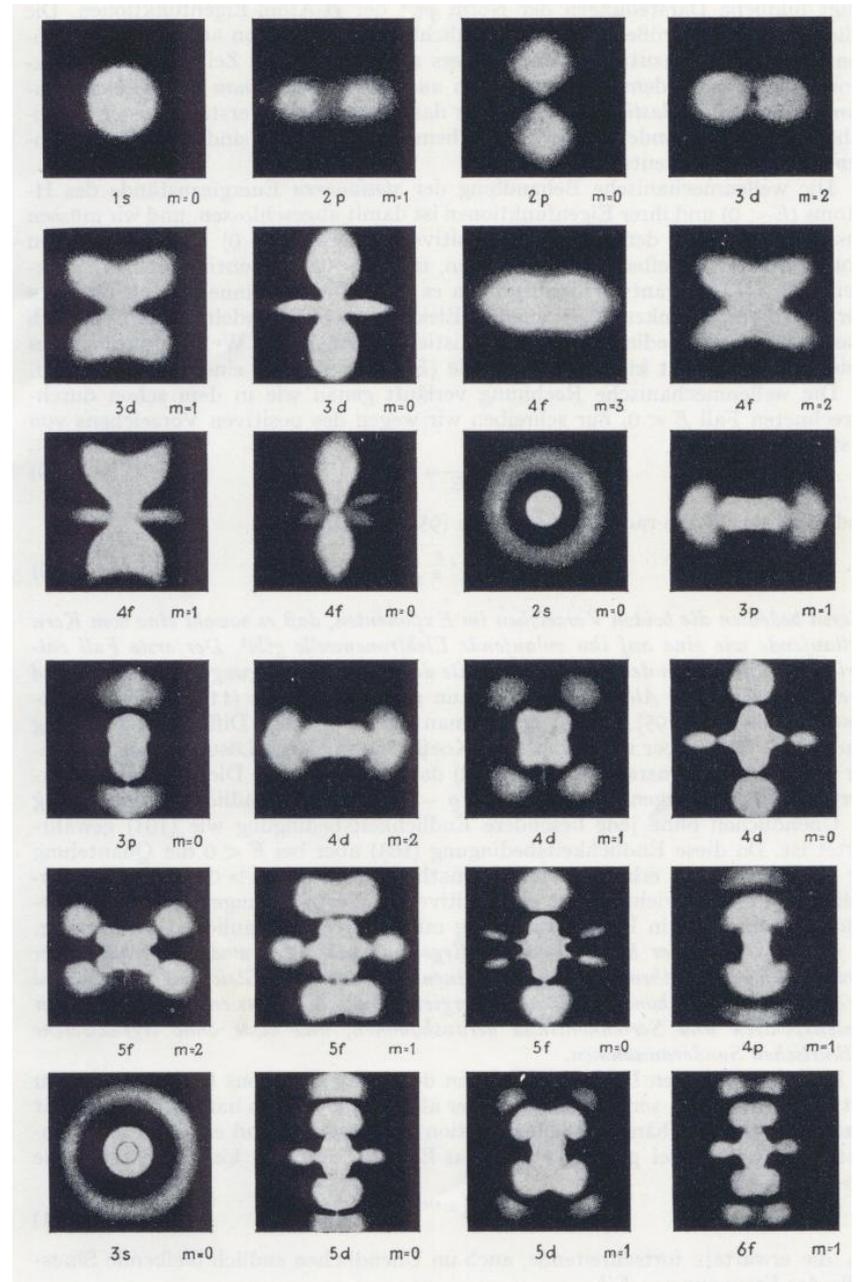
Bohr's Theory



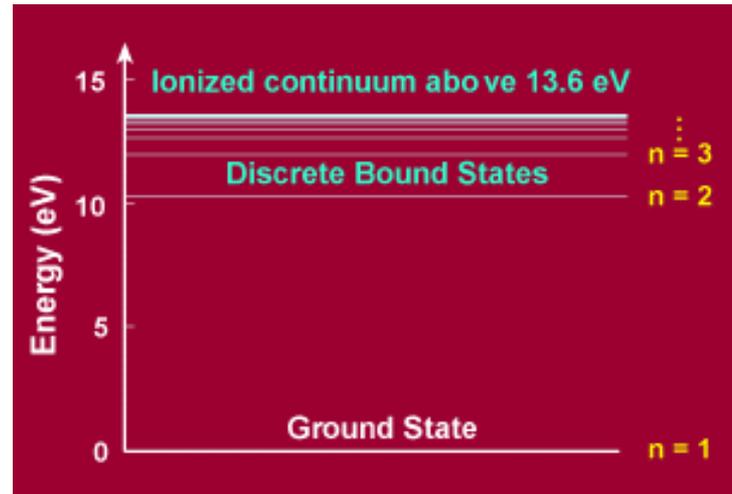
Structure of Hydrogen Atoms

Quantum Mechanics

Schrödinger & Bohr Atomic Models



E1-Example 3



Calculate the five lowest energy states including the ground state of hydrogen atoms in eV.

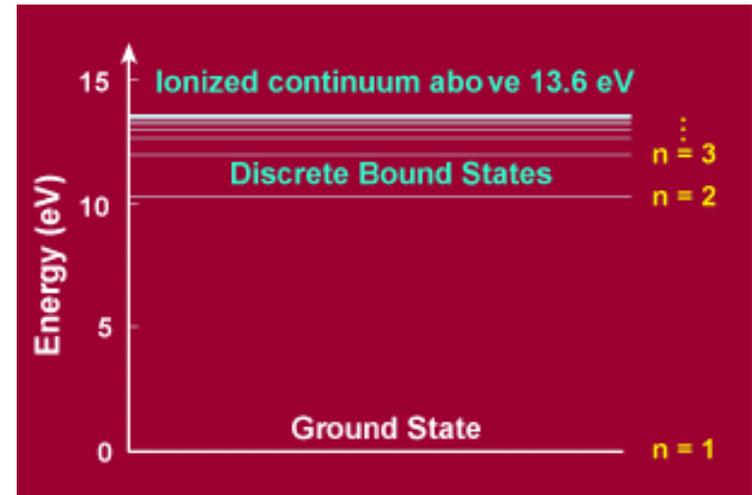
Calculate the ionization potential of hydrogen atoms in eV.

Find the energy states of hydrogen atoms that can absorb photons with wavelengths in the visible range (400-700 nm).

Calculate the five lowest energy states including the ground state of hydrogen atoms in eV.

$$\Delta E = R_H \left[\frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$$

$$E_n = 13.5992 \left(1 - \frac{1}{n^2} \right) \quad [eV]$$



$$n = 1 \quad E_1 = 0 \quad [eV]$$

$$n = 2 \quad E_2 = 10.2$$

$$n = 3 \quad E_3 = 12.1$$

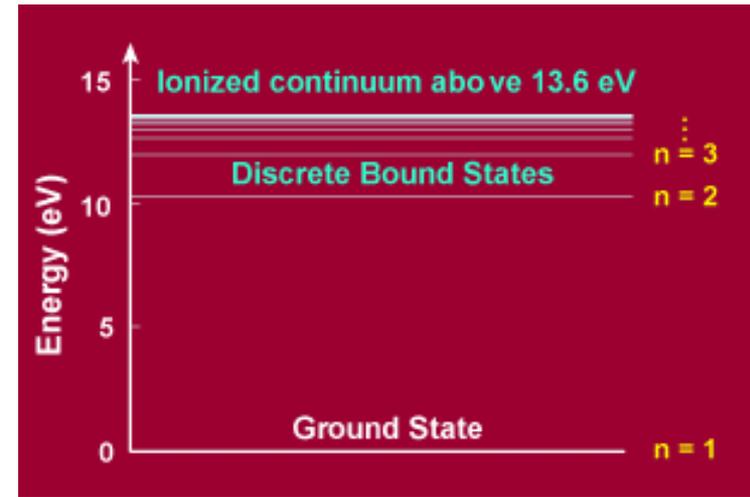
$$n = 4 \quad E_4 = 12.7$$

$$n = 5 \quad E_5 = 13.1$$

Calculate the ionization potential of hydrogen atoms in eV.

$$E_n = 13.5992 \left(1 - \frac{1}{n^2} \right) \quad [eV]$$

$$n = \infty \quad E_{ionization} = 13.5992 \quad [eV]$$



Find the energy states of hydrogen atoms that can absorb photons with wavelengths in the visible range (400-700 nm).

$$E = h\nu = \frac{hc}{\lambda} = \frac{6.626 \times 10^{-34} \times 3 \times 10^8}{\lambda \times 10^{-9} \text{ [nm]}} = \frac{1.9878 \times 10^{-16}}{\lambda} \quad [J] = \frac{1240}{\lambda} \quad [eV]$$

$$\lambda = 400 \text{ nm} \quad \Rightarrow \quad 3 \quad [eV]$$

$$\lambda = 700 \text{ nm} \quad \Rightarrow \quad 1.8 \quad [eV]$$

$$n = 1 \quad E_1 = 0 \quad [eV]$$

$$n = 2 \quad E_2 = 10.2$$

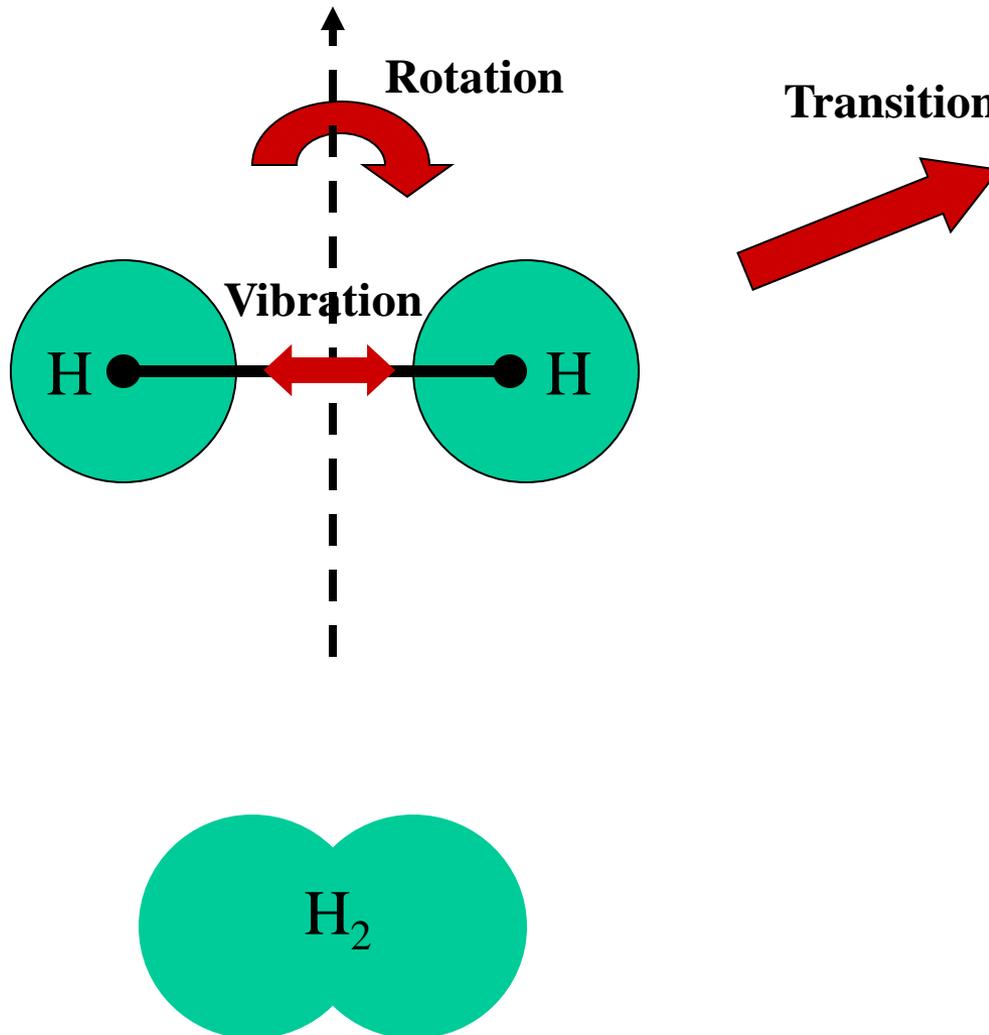
$$n = 3 \quad E_3 = 12.1$$

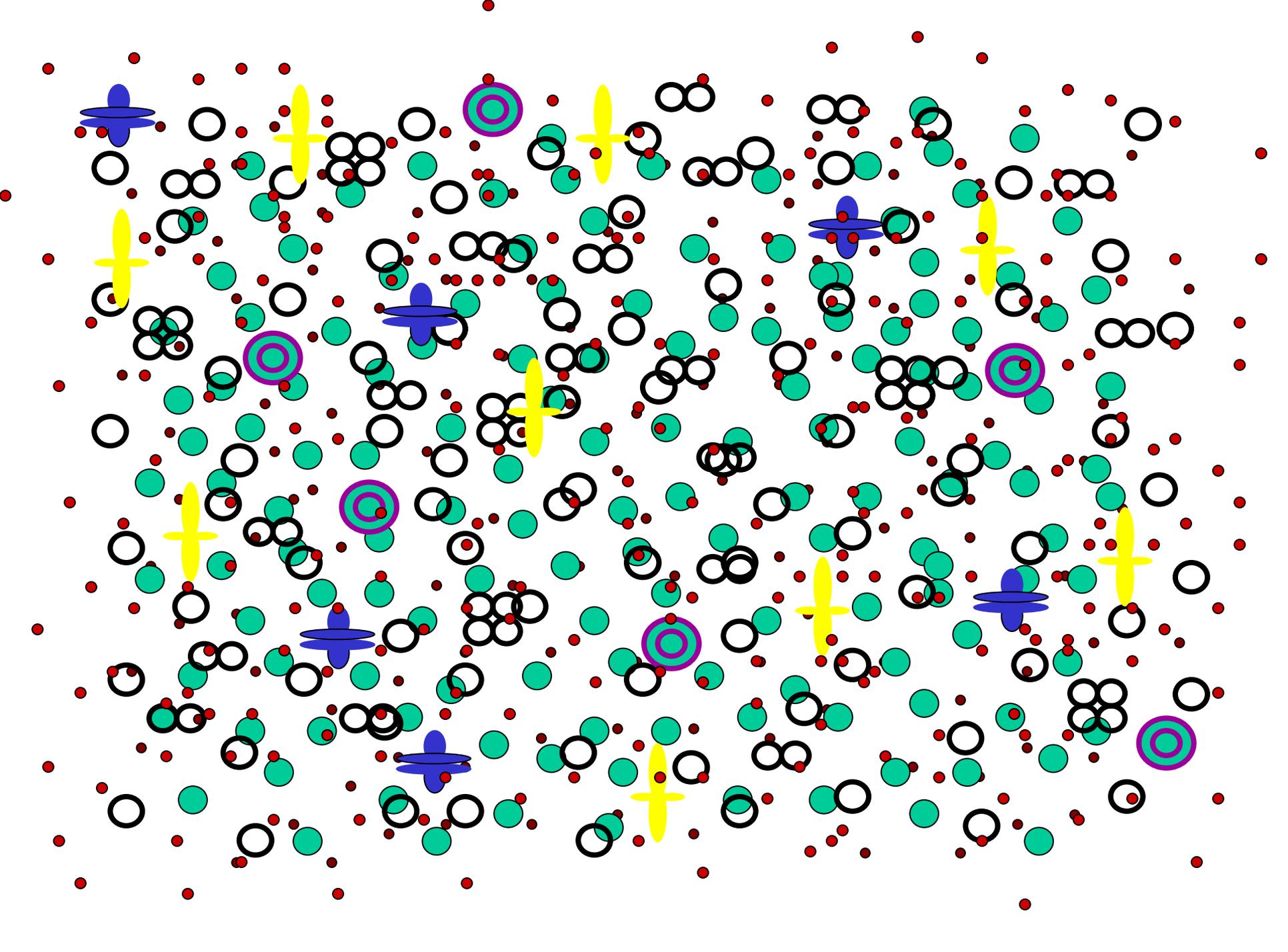
$$n = 4 \quad E_4 = 12.7$$

$$n = 5 \quad E_5 = 13.1$$

$$n_{upper} = 3, 4, 5, \dots \quad \rightarrow \quad n = 2$$

Hydrogen Molecule





What is pressure of an ideal gas?

Number of particles per m⁻³

$k=1.38 \times 10^{-23}$ J/K Boltzmann Constant

Pressure

$$p = n \times k \times T$$

Temperature

[Pascal=N/m²; 1 atm=760 Torr =101325 Pa]

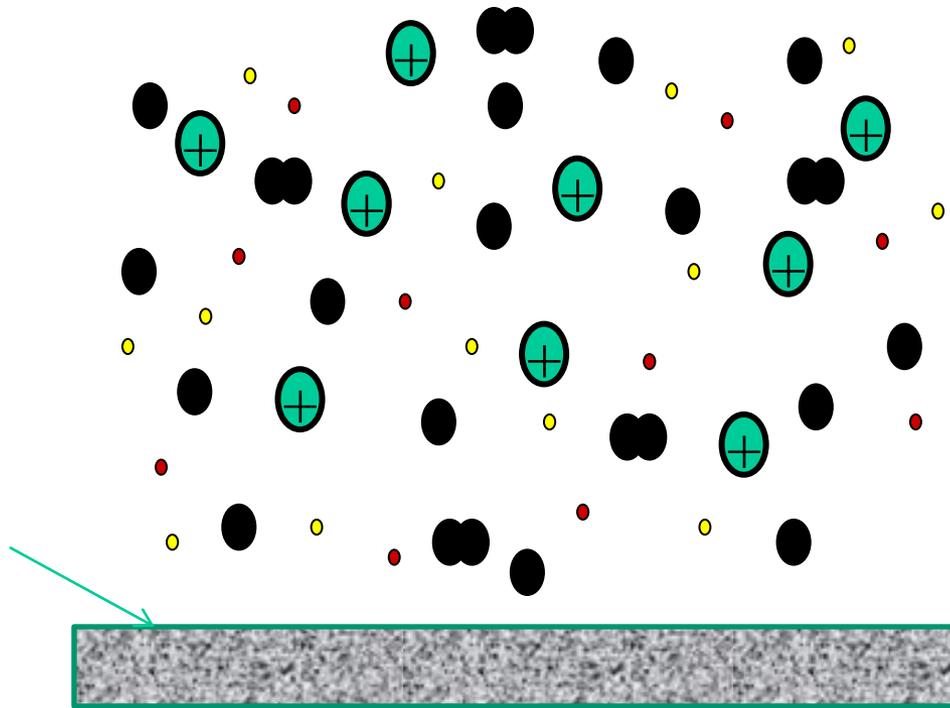
$$p = n_e k T_e + n_{ion} k T_{ion} + n_a k T_a$$

Electron
Partial Pressure

ions
Partial Pressure

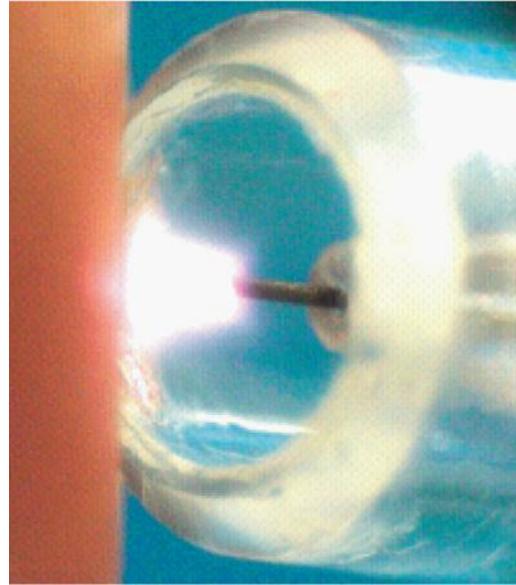
Neutral species
Partial Pressure

Total number of collision on the substrate per unit area per unit time



$$= \frac{1}{4} n \bar{v} = \frac{1}{4} n_e \bar{v}_e + \frac{1}{4} n_{ion} \bar{v}_{ion} + \frac{1}{4} n_a \bar{v}_a$$

Problem 2



How many electrons, He and He⁺ are impacted on the surface per unit area and per second. Number density of He and electrons are approximately 10^{17} and 10^{13} [m⁻³], respectively.

Number of collisions on the substrate
Per unit area and per unit time, n_c . $= \frac{1}{4} n \bar{v}$

$$n_{c,He} = \frac{1}{4} \times 10^{17} \times 1.24 \times 10^3 = 3.1 \times 10^{20}$$

$$n_{c,He^+} = \frac{1}{4} \times 10^{13} \times 1.24 \times 10^3 = 3.1 \times 10^{16}$$

$$n_{c,e} = \frac{1}{4} \times 10^{13} \times 9.47 \times 10^5 = 2.4 \times 10^{19}$$

Projects